

SPECTRUM OF THE FLAME ABOVE A COPPER ARC IN AIR

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Plate III

ABSTRACT. The spectrum of the flame above a copper arc in air has been photographed on a glass prism spectrograph with a dispersion of about 12 Å per mm at $\lambda 5500\text{Å}$, and all the radiations between $\lambda 4100$ and $\lambda 5400$ Å have been measured. A large number of bands exist. The origin and structure of most of them do not seem to be very well understood. A tentative arrangement for three systems of bands is given.

INTRODUCTION

The spectrum of the flame above a copper arc in air does not seem to have received the attention it deserves, by virtue of being associated with such a common source of radiation in a spectroscopic laboratory. It is believed that the radiation of the flame is chiefly due to CuO which has an intense system of band in the orange and red region and some weak bands in the green. Several measurements of the orange and red bands have been published (Mulhken, 1925; Mahanti, 1930; Loomis and Watson, 1935; Pearse and Gaydon, 1941; Rosen, 1945 and Guntzsch, 1946), and some of these workers have also attempted an analysis of them, each in his own way. As for the bands in the other regions there does not seem to exist even a complete measurement of them (Hertenstein, 1912, Pearse and Gaydon, 1941; Rosen 1945; Lejune and Rosen, 1945 and Rao, 1946). It was, therefore, considered necessary to photograph the spectrum of this flame and to make a complete measurement of the wavelengths of all the radiations present in it. This has been done and as a result it has been found that the spectrum is quite complicated and the structures are not easily understood. Full details are given in the following sections:

EXPERIMENTAL

The copper arc used was run from 110 volts D. C. supply. The electrodes were horizontal and the image of the flame above the level of the electrodes was thrown by means of a lens over the slit of a glass Littrow spectrograph having a dispersion of about 12 Å per mm. at $\lambda 5500$ Å. Exposures of about one and a half hours on Ilford Special Rapid Panchromatic plates were necessary to get the fainter bands with an intensity so that they could be read in the microscope. The iron arc was used for comparison. A copper arc was also exposed side by side in order to eliminate the copper lines from the spectrum of the flame.

DESCRIPTION OF THE SPECTRUM

The spectrum of the flame is given in Plate III. The prominent features of the spectrum are the very intense bands in the orange region, one or two intense bands in the green, a large number of comparatively faint bands in the orange, yellow and green and the strong green and yellow lines of copper, which could not be eliminated, although their relative intensities have been considerably reduced. The fainter bands can apparently be put into several groups described below.

Beginning from the short wavelength side, the first group starts at $\lambda 4828\text{\AA}$ and extends up to $\lambda 4916\text{\AA}$. There are 9 prominent bands in this region with heads at $\lambda 4828.5$, 4836.5 , 4845 , 4852.5 , 4862.5 , 4871.5 , 4882.5 , 4893.5 and 4916\AA . The last band is the most intense of the lot, although the others also are quite strong. The fainter bands of this group have heads at $\lambda 4852.5$, 4876 , 4901 , 4904 and 4911\AA .

The second group lies between $\lambda 4976$ and $\lambda 5040\text{\AA}$. Two bands at $\lambda 5030.3$ and 5039.5\AA have distinct heads, both degraded to the red. The bands at $\lambda 4976.1$, 4982.7 and 4989.6\AA are somewhat less intense. These are also degraded to the red. There is a strong band at $\lambda 4996\text{\AA}$. This is quite broad and does not show any head. The measurement corresponds to the centre of the band. It is possible that there may be a faint structure at $\lambda 5002\text{\AA}$ which is masked by the above band. The other faint structures are at $\lambda 5005.9$, 5008.7 and 5011.9\AA .

A single intense band lies at $\lambda 5071.4\text{\AA}$. It is degraded to the red but is quite blurred. Another intense band occurs with head at $\lambda 5227.8\text{\AA}$, degraded to the red.

A third group starts at $\lambda 5274\text{\AA}$, 4 distinct bands appear with heads at $\lambda 5274.0$, 5278.9 , 5282.7 and 5285.6\AA . They all appear degraded to the red, but they are all so narrow that they look more or less like lines. There are faint band heads at $\lambda 5302.6$ and 5304.6\AA , followed by two stronger ones at $\lambda 5307.4$ and 5312.0\AA .

Again there are two bands, comparatively isolated, the intense one with head at $\lambda 5345.1\text{\AA}$ and the faint one, at $\lambda 5347.9\text{\AA}$.

A fourth group of bands appears between $\lambda 5450$ and $\lambda 5500\text{\AA}$. Nine heads have been measured, which seem to be gradually converging towards the short wavelength side. They all appear degraded to the red. The more prominent amongst them have their heads at $\lambda 5496.1$ and 5488.2\AA . The others follow at shorter wavelengths with gradually decreasing separation and intensity.

A moderately intense band is observed at $\lambda 5529.5\text{\AA}$, degraded towards the red. Another one which is a little blurred appears at about $\lambda 5557\text{\AA}$.

A fifth group of bands is observed between $\lambda 5827$ and $\lambda 5895\text{\AA}$. Some of the bands of this group are blurred. The stronger and more distinct ones are at $\lambda 5847.6$ and 5856.0\AA .

TABLE I

Wavelengths of bands in the flame above a copper arc in air.

$\lambda_{\text{air}}(\text{\AA})$	$\nu_{\text{vac}}(\text{cm}^{-1})$	Int.	$\lambda_{\text{air}}(\text{\AA})$	$\nu_{\text{vac}}(\text{cm}^{-1})$	Int.
4628.4	20704.5	s	5833.3	17138.2	m
36.5	670.4	s	41.9	111.0	w
45.2	633.2	s	43.4	108.6	
52.6	601.8	s	47.6	96.3	m
54.9	592.0	m	50.4	98.1	?
62.4	560.2	s	56.0	97.8	m
65.7	546.3	w	59.0	93.0	w
71.4	522.3	s	66.0	94.7	m
75.8	503.7	m	73.7	92.3	w
82.6	475.2	s	80.3	90.2	w
89.1	448.0	w	89.1	16975.8	m
93.5	420.6	s	95.0	95.8	m
4901.0	398.5	w	5932.6	879.8	w
04.2	385.0	m	26.4	890.0	w
10.9	357.2	m	32.2	852.5	m
16.2	335.3	vs	36.0	841.7	w
76.1	090.5	w	40.5	828.9	m
82.7	063.9	m	43.6	820.2	vw
89.6	036.1	m	48.2	807.2	m
95.7	011.6	s	50.5	800.7	w
5007.6	19984.0	w	57.5	780.9	m
05.9	970.9	w	65.0	759.8	m
08.7	959.7	w	0030.1	578.1	?
11.0	947.0	m	45.1	537.8	vv's
30.3	874.0	s	46.0	535.3	?
39.5	837.7	s	48.3	529.0	s
71.4	712.0	vs	52.7	517.0	s
5227.8	123.2	vs	59.1	499.5	vv's
32.4	106.4	m	64.1	485.9	w
34.3	090.4	s	69.0	172.6	w
35.8	094.0	s	78.2	147.7	w
46.5	091.4		0146.8	264.1	vv's
41.8	072.1	vv's	51.1	252.8	w
74.0	18955.7	s	57.3	236.4	m
78.0	938.1	s	61.5	225.3	vv's
82.7	924.5	s	68.7	206.4	m
85.6	914.1	w	74.5	191.2	w
92.6	889.7	m	91.5	146.7	vw
5302.6	853.4	?	6205.3	110.8	w
04.6	846.3	w	11.6	094.5	w
07.4	836.4	m	20.1	072.5	m
12.0	820.1	vs	23.0	065.0	m
45.1	703.5	s	70.3	15943.8	w
47.9	693.7	?	80.1	918.9	w
5385.0	18565.0	?	89.1	806.1	w
5457.7	317.7	w	94.0	883.8	vv's
66.1	289.5	w	97.6	874.7	s
68.7	280.8	w	0302.6	862.1	m
72.7	267.5	w	13.5	834.7	m
77.0	253.1	m	23.9	808.7	w
82.2	235.8	m	78.5	074.3	w
88.2	215.9	m	84.9	657.6	vw
96.1	189.7	m	92.5	630.1	vw
5529.5	079.8	s	0400.3	620.0	s
56.8	17991.0	m	03.5	612.2	vw
57.9	987.4	s	15.2	583.7	vw
5827.6	155.0		30.2	547.3	vw
29.0	150.9	m			

A sixth group of bands appears between $\lambda 5922.6$ and $\lambda 5965\text{\AA}$. This group of bands is followed by the very strong bands in the orange and red regions extending up to $\lambda 6550\text{\AA}$. Our measurements in this region are in general agreement with those due to Pearse and Gaydon (1941). The bands are degraded to the red. The structure is complex, consisting of several heads of varying intensity.

A complete list of all wavelengths measured is given in Table I, and can be generally taken to be correct to within $\pm 1\text{\AA}$.

The intensities given are visual estimates, and the notations in that column have the following significance: *s* means strong, *m* means medium, *w* means weak, *v* means very and ? means that the band measured was so faint that its existence is doubtful. Measurements included in braces mean that the corresponding bands did not clearly indicate which way they were degraded and in such cases either both the ends and the centre or at least both of them have been measured. All the other bands were degraded to the red.

DISCUSSION

We will consider the strong bands in the red and orange and the fainter ones dispersed over the entire yellow and green regions separately.

(a) *Bands in the orange and red regions.* An analysis of the bands in this region was proposed by Mahanti (1930), but this is doubted by Loomis and Watson (1935). Rosen (1945) carried out a vibrational analysis of these bands and found that they could be represented by the equation,

$$\nu = 16222 - 625 \nu'' + 3 \nu''^2 + 274 \nu'$$

A different equation for the vibrational analysis of these bands has been proposed by Guntch (1946). According to him

$$\nu_0 = 16273.9 + 285.6 u' - 9.2 u'^2 - 318.6 u'' + 4.4 u''^2, \text{ where } u = v + \frac{1}{2}$$

Guntch gives evidences in support of his equation. He has also carried out the rotational analysis for 5 bands of this system and believes them to be due to $^2\Sigma - ^2\Sigma$ transition of CuO molecule. The *B* values are $B''_0 = 0.632$ and $B'_0 = 0.591 \text{ cm}^{-1}$. Our own arrangements, as given in Table II, is in agreement with Guntch's equation, but it suffers from certain defects, viz, (1) the agreement is not as good as could be expected, (2) the vibrational differences are not very satisfactory and (3) all the bands in the region are not accounted for by this arrangement.

(b) *Bands in the green and yellow regions.* Although some of the measurements given in Table I in the yellow region may be associated with the orange bands, it seems doubtful if all of them can be accounted for in this way. Nor does there appear any clue for classifying the yellow and the green bands from an inspection of the spectrum. At any rate, then

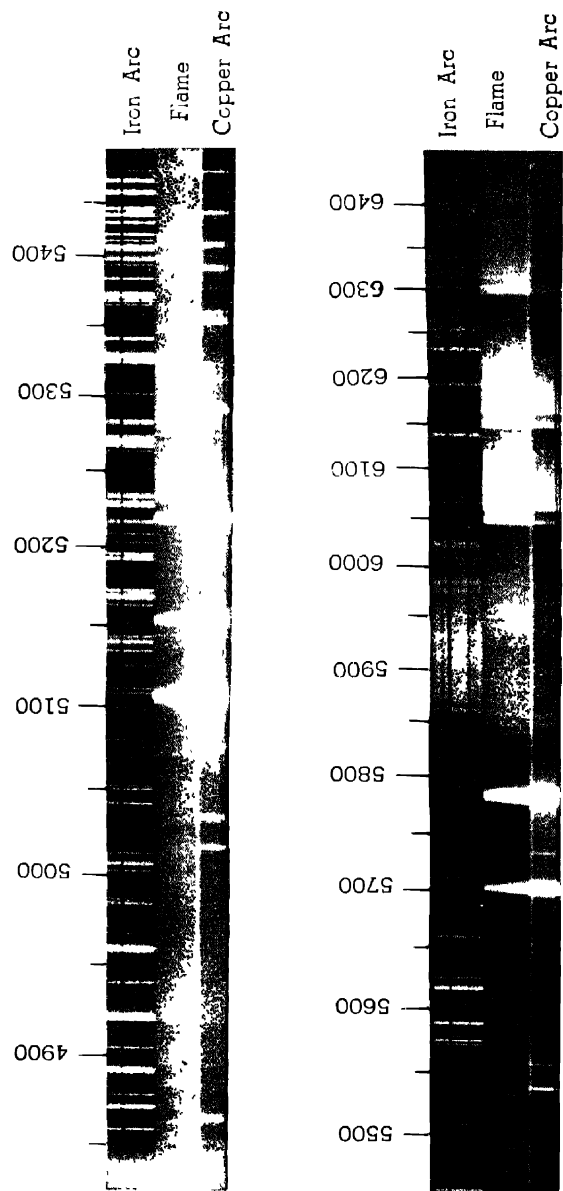


Fig 1.—Spectrum of the flame above a Copper arc in air

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appearance is so different from that of the orange and red bands that it does not look probable that the entire stuff can be attributed to the same emitter. Yet certain bands, from their appearance, seem to suggest some sort of regularity. Thus the bands, referred to as belonging to the first group in the previous section, indicate, from their appearance, that it should be possible to arrange them in some sort of fragmentary system. This is easily done, as given in Table III.

TABLE II
 $v' - v''$ scheme for orange-red bands

$v' \backslash v''$	0	1	2	3	4	5
0	16264					
1	16538	6225	15920	15670	15310*	
2		16499		15884		15270*
						15547

* Denotes that the corresponding band has been taken from Pearse and Gaydon (1941).

TABLE III
 $v' - v''$ scheme for the bands between $\lambda 4828$ and $\lambda 4916\text{\AA}$.

$v' \backslash v''$	0	1	2	3
0	20504	20385		
1	20592	20475	20357	
2	20670	20560	20448	20335
3		20633		20430

The ω 's are, however, so much different from those suggested for the orange-red system by Guntch that it is difficult to believe that those two systems are due to the same emitter.

The bands in the other groups mentioned in the previous section do not generally seem to suggest any order. Some bands, however, between $\lambda 5270$ and $\lambda 5530\text{\AA}$, can be picked up from the spectrogram and their separation indicated that they might belong to one system. Such bands are given in Table IV. This system may be attributed to the same emitter to which the orange system belongs, *i.e.*, to CuO. But here again none of the ω 's agree and if both the systems belong to CuO, the states involved are all different.

TABLE IV

 $\nu' - \nu''$ scheme for some bands in the green region.

$\nu' \backslash \nu''$	0	1	2	3	4
0	19093*	18827	18565	18318	18080
1		18956	18704		

* This band is strong, but its head is not distinct. The centre had to be measured. More work seems necessary to clarify all these points.

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